Developing A Universal Radio Astronomy Backend Dr. Ewan Barr, MPIfR Backend Development Group

Overview

- Why is it needed?
- What should it do?
- Key concepts and technologies
- Case studies:
 - MeerKAT FBF and APSUSE instruments
 - EDD, TNRT and SKA Prototype systems

Why is a URAB needed?

Adaptability:

- Rapid development and deployment of new (and old) processing algorithms
- Adapt to meet new needs (e.g. real-time transient detection using ML)

Commensality:

The ability to observe for multiple distinct disciplines simultaneously

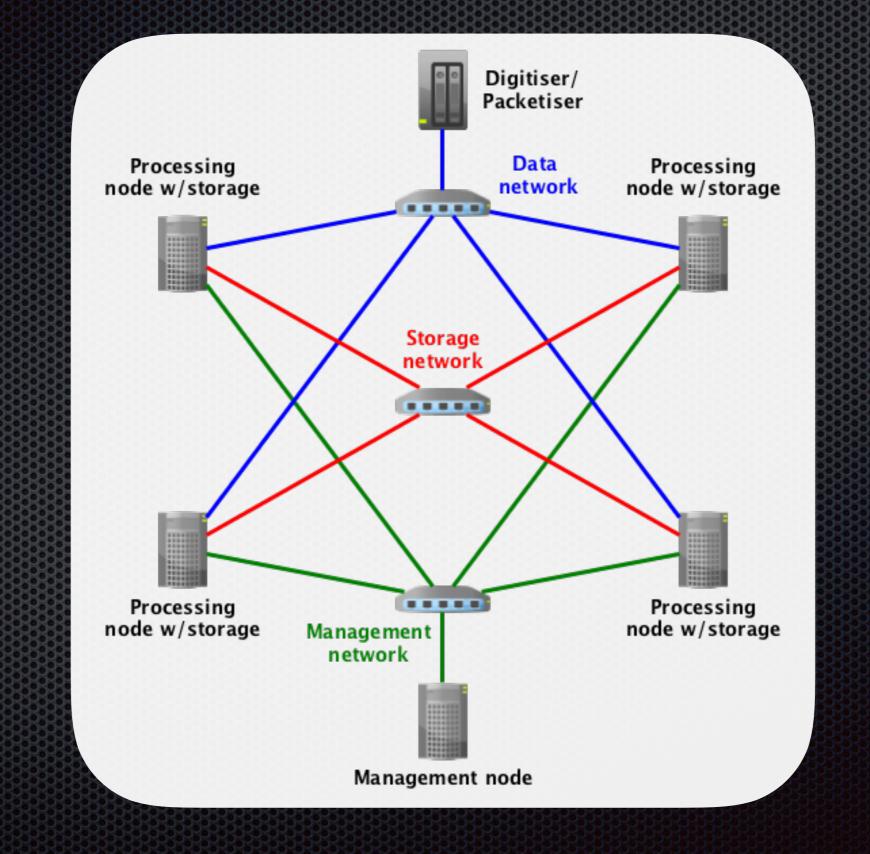
Simplification:

The move from custom hardware to COTS-based systems opens the developer pool lowering the cost of development

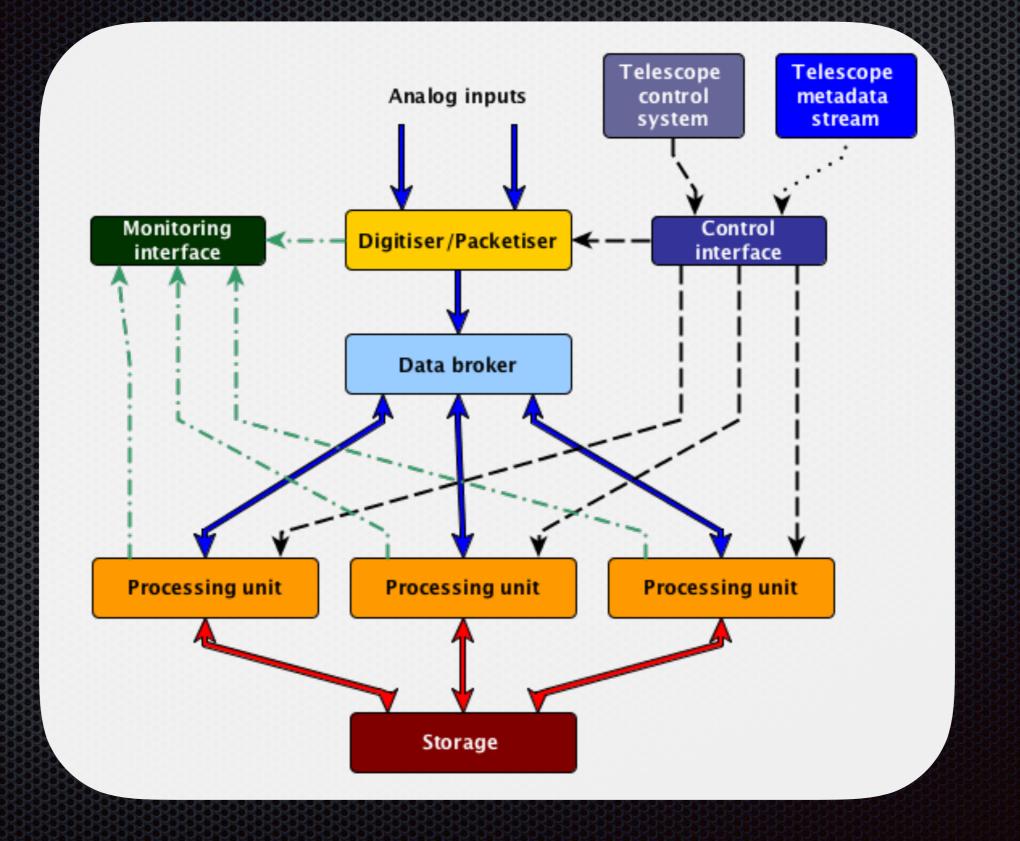
What should it do?

- Satisfy basic and not so basic telescope processing needs:
 - Digitisation, channelisation (Hz MHz), Stokes detection, integration
 - VLBI, pulsar timing, real-time transient search
 - Dumpable voltage buffers, online RFI flagging (e.g. SK)
- Produce standard-format science-ready outputs (FITS, VDIF, FIL, etc.)
- Run out-of-the box (no installation necessary)
- Provide rich feedback to operators and astronomers

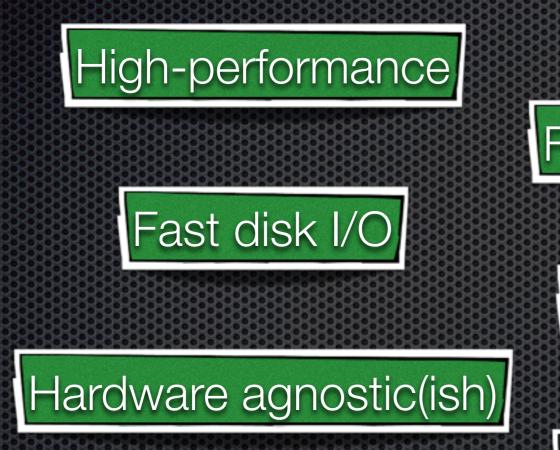
Physical view



Functional view



Requirements





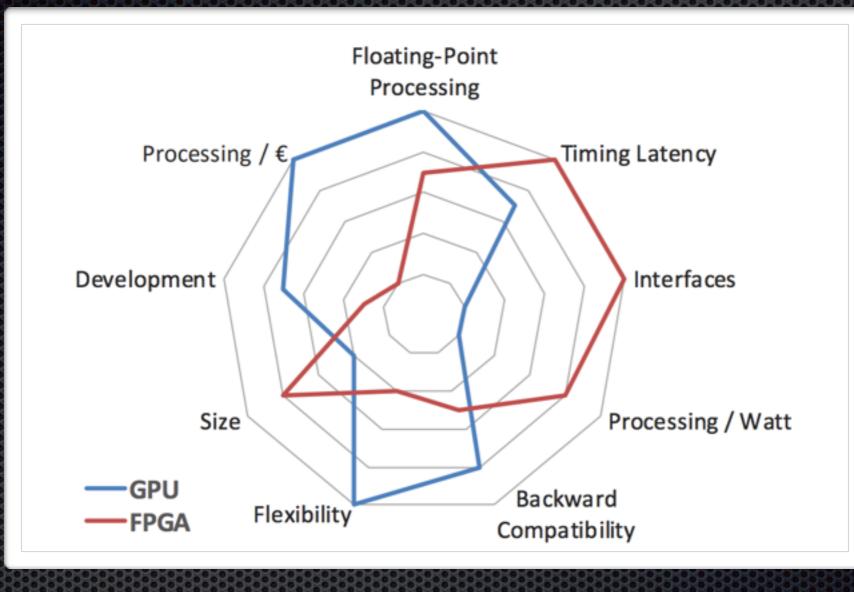
Flexible data transport

Monitoring & Reporting

Standardised interfaces

High-performance

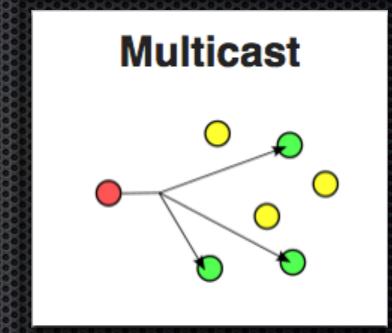
Credit: BERTEN DSP

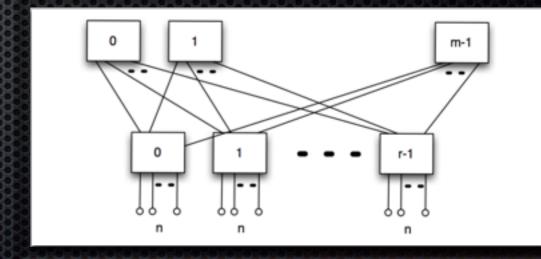


- GPUs are the preferred due to flexibility, cost and FP32 performance
- PCIe mounted FPGAs are interesting possibility due to native 100 GbE support (e.g. Nallatech 520N)

Flexible data transport

- Ethernet data backbone
- Data streams split across multiple groups
- Low traffic per group (6 Gb/s)
- Highly scalable
- Self load balancing





Standardised interfaces

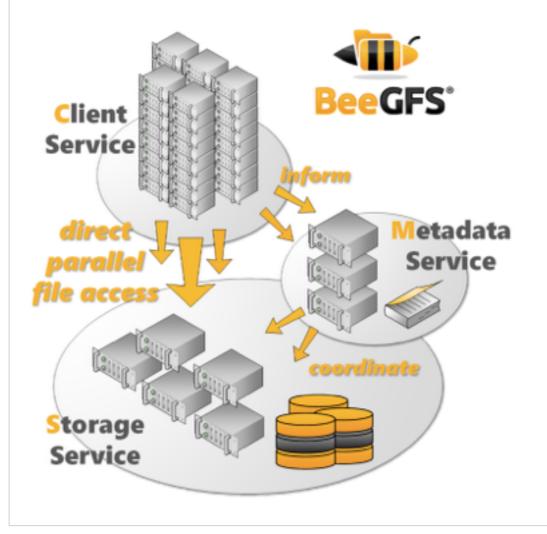
- Network data: VDIF, SPEAD
- Control interface: Tango, KATCP
- Application data: PSRDADA, HASHPIPE
- Metadata: KATCP, redis, etcd



- Processing nodes are also storage nodes
- Performance increases with number of spindles
- Infiniband or Ethernet (w/ RoCE) interconnect
- Cheap storage nodes can be added to improve performance

BeeGFS

The Parallel Cluster File System



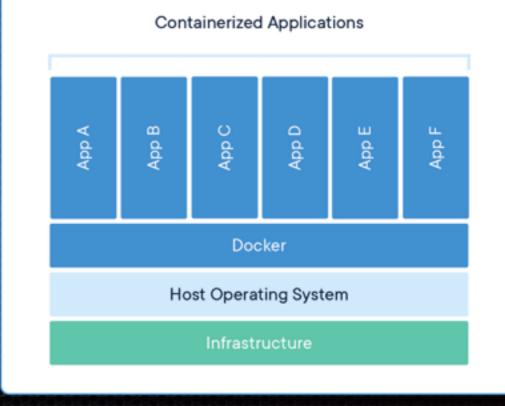
Hardware agnostic(ish)

- Containerisation
- Resource virtualisation
- Version control
- Environment control



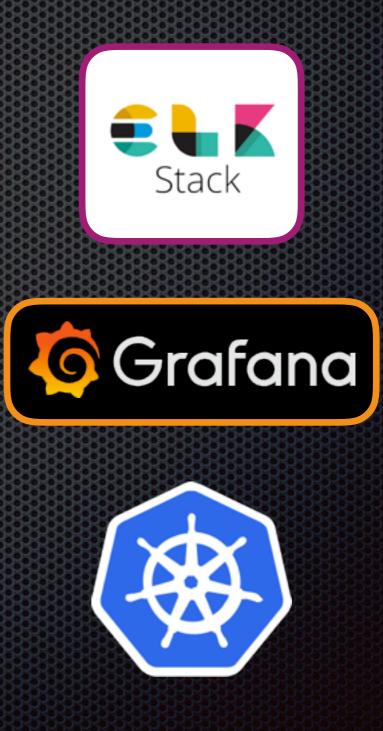
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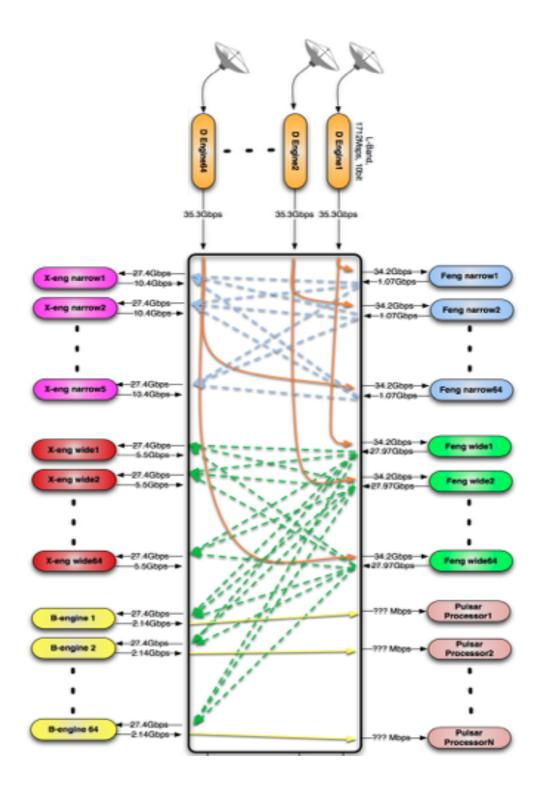


Monitoring & Reporting

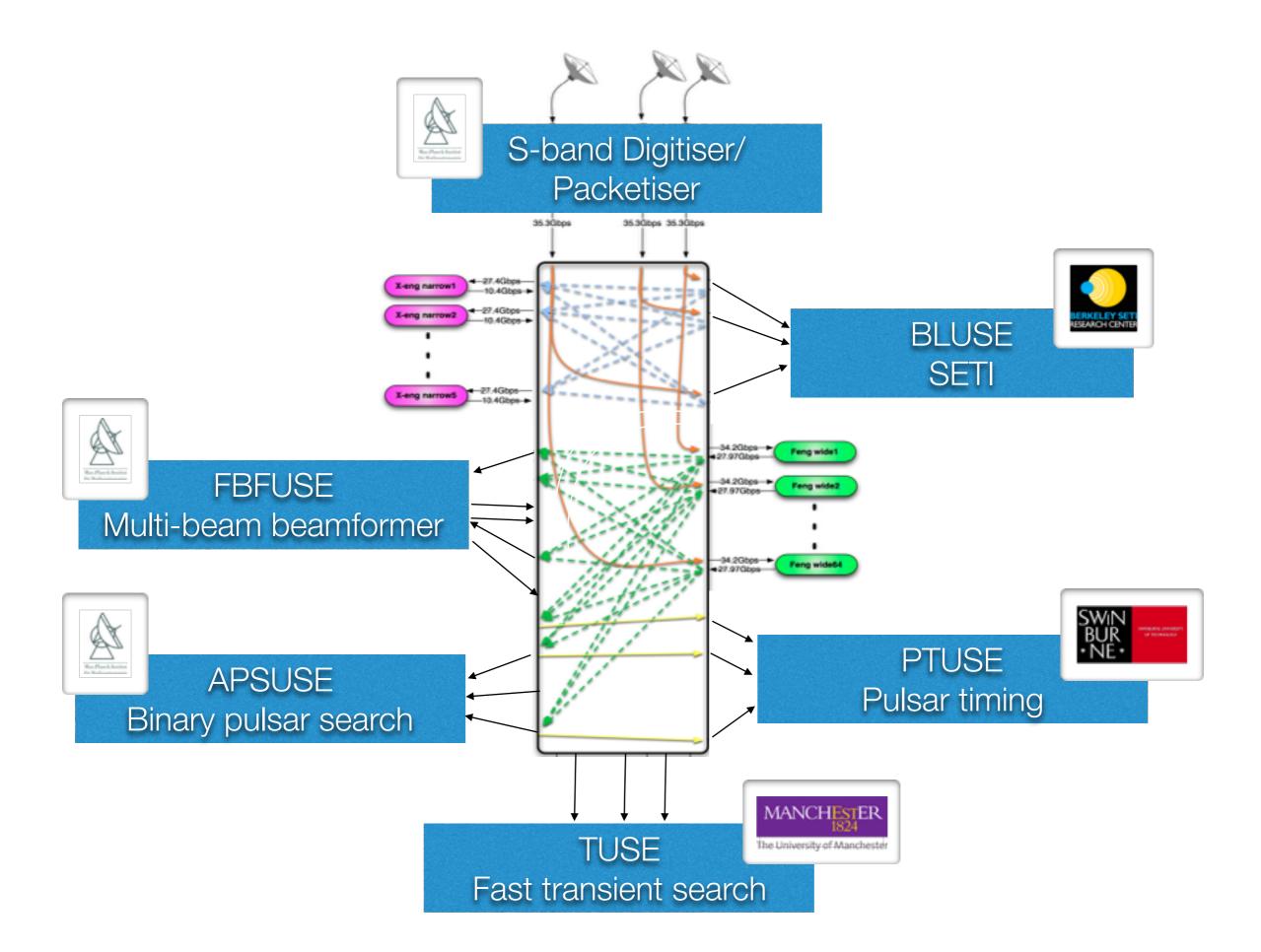
- Unified logging: Elasticsearch, Logstash, Kibana
- Hardware monitoring: Grafana,
 Collectd, Prometheus,
 Heapster
- Everything run as services using Kubernetes
- Application monitoring: ???



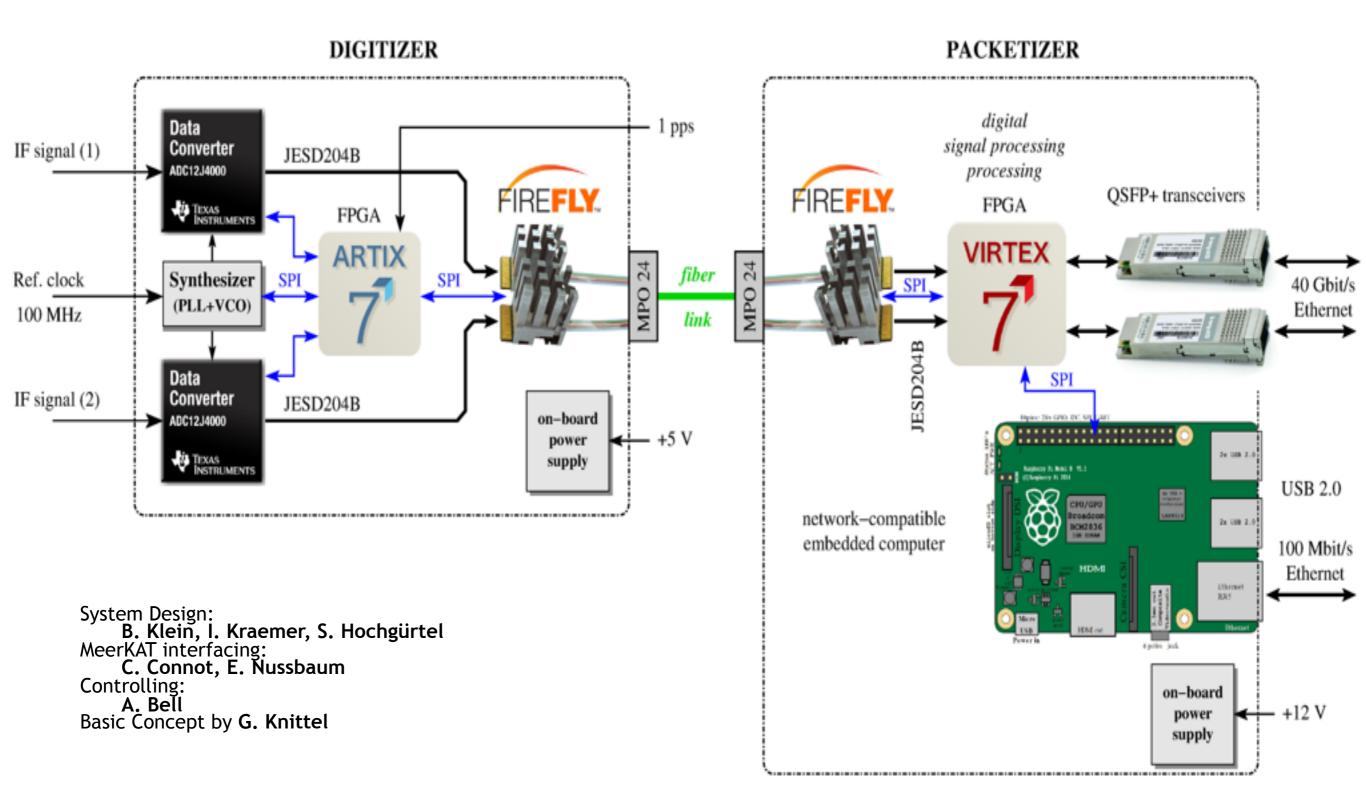
Case study I: MeerKAT

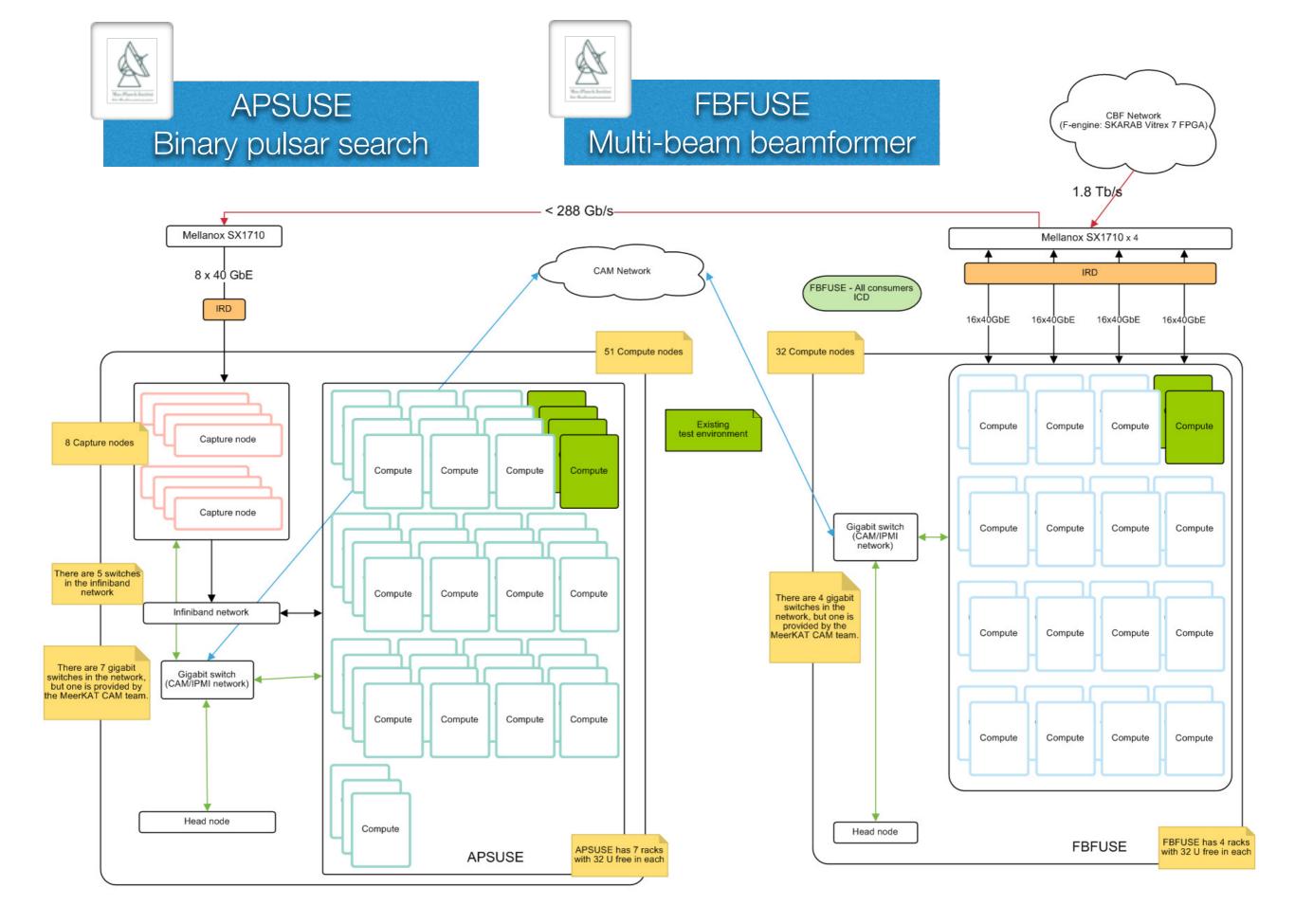


MeerKAT CBF switch







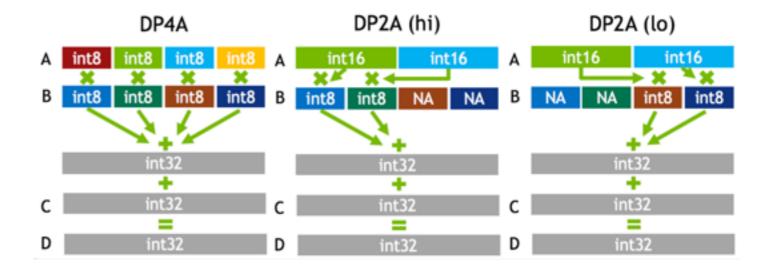


FBFUSE Cluster Specifications

Hardware – Compute Node (32x)	Huawei FusionServer 2288H V5 2x Xeon Gold 6134 2x 40 GbE NIC 384 GB RAM (transient buffer) 2x GTX 1080 Ti				
Performance	1 Petaop				
Transient buffer	12 TB (~50 seconds)				

FBFUSE – BENCHMARKING

 DP4A support gives 4x
 performance over f32 (4 Tflops to 16 Tops on Titan X Pascal)



# Antennas	Beamformer benchmarking (Nbeams 85% real-time)						
	856 MHz	428 MHz	214 MHz				
4	2944	5888	11776				
8	3008	6016	12032				
16	2272	4544	9088 7040 3840				
32	1760	3520					
64	960	1920					

Original prototype: https://github.com/ewanbarr/beanfarmer Integrated version: https://github.com/ewanbarr/psrdada_cpp



APSUSE Cluster Specifications

Hardware – Capture Node (8x)	Huawei FusionServer 2288H V5 2x Xeon Gold 6136 40 GbE & 56 Gb/s IB NICs 192 GB RAM			
Hardware – Compute Node (60x)	Huawei FusionServer 2288H V5 2x Xeon Silver 4116 56 Gb/s IB NIC 96 GB RAM 2x GTX 1080 Ti			
Storage volume	3.5 PB (distributed)			

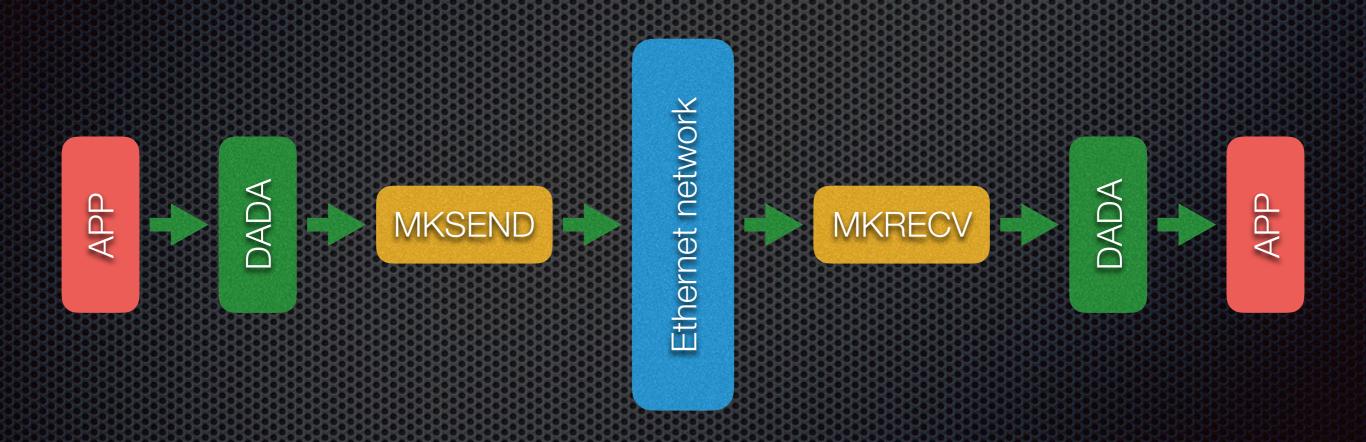
Write speed

48 GB/s (sustained)

MKSEND / MKRECV

- Based on SPEAD2 library:
 - https://casper.berkeley.edu/wiki/SPEAD
 - https://github.com/ska-sa/spead2
- MKRECV: Captures SPEAD stream(s) and saves to PSRDADA buffers
- MKSEND: Converts PSRDADA buffer to SPEAD stream(s)
- ASCII configuration file to support arbitrary SPEAD streams
- Uses Infiniband Verbs for kernel bypass

MKSEND / MKRECV



MKRECV configuration

Network connection PACKET_SIZE 1500 IBV_IF 192.168.2.20 PORT 7148 MCAST_SOURCES 224.2.1.150,224.2.1.151,224.2.1.152,224.2.1.153 DADA_KEY dada SYNC_TIME 1231235243.0000000 SAMPLE_CLOCK 1750000000.0 NTHREADS 32

#MeerKat F-Engine NINDICES 3

The first index item is the running timestamp
IDX1_ITEM 0
IDX1_STEP 2097152 # The difference between successive timestamps

This second index is the F-engine ID IDX2_ITEM 1 IDX2_LIST 0,1,2,3,5,6,7,8,9,10,11,12,13,14,15 # Antennas to capture

The second index item is the frequency IDX3_ITEM 2 IDX3_LIST 0,256,512,768,1024,1280 # List of frequency partitions

Case study II:







Eifelsberg Direct Digitisation Backend (EDD)

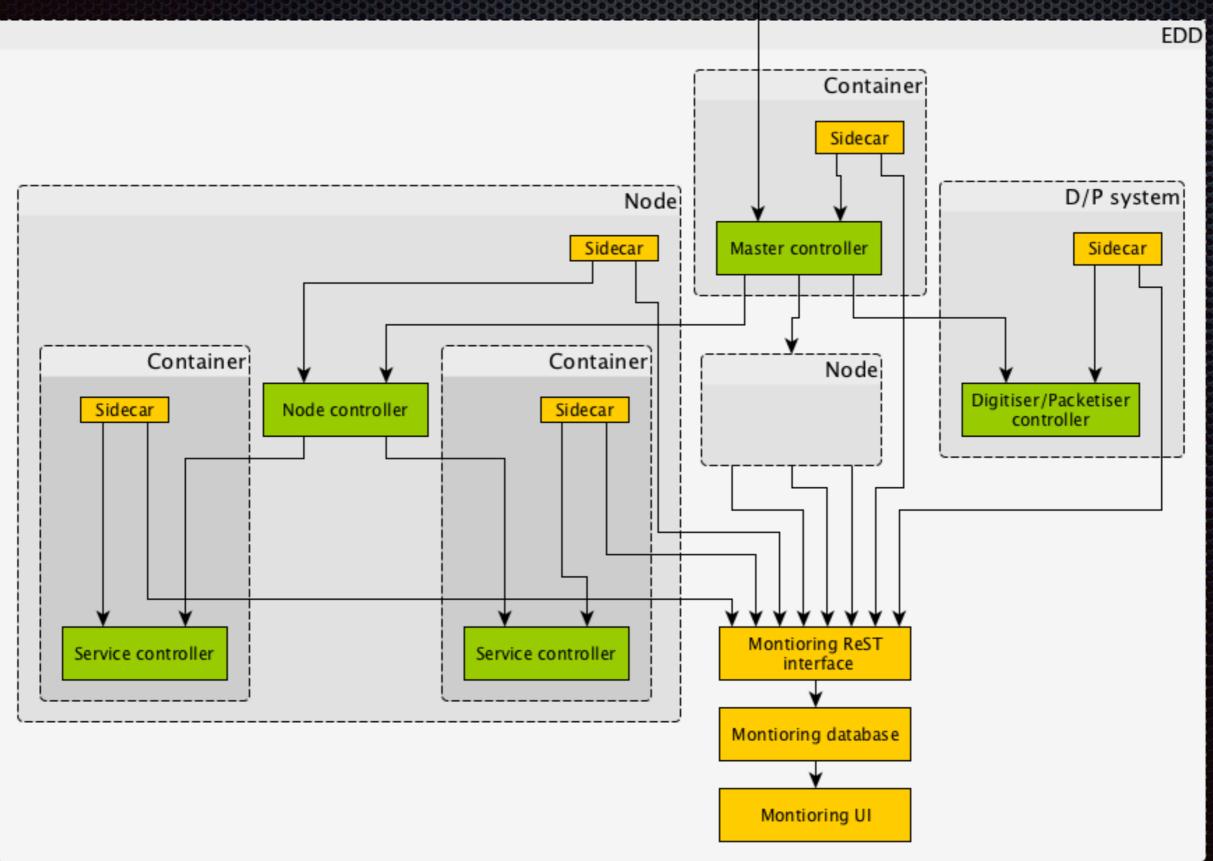
- The Effelsberg realisation of the Universal Backend concept
- Intended to support the new range of direct digitisation receivers:

K, C+, UBB, Q, Ka and Ku bands

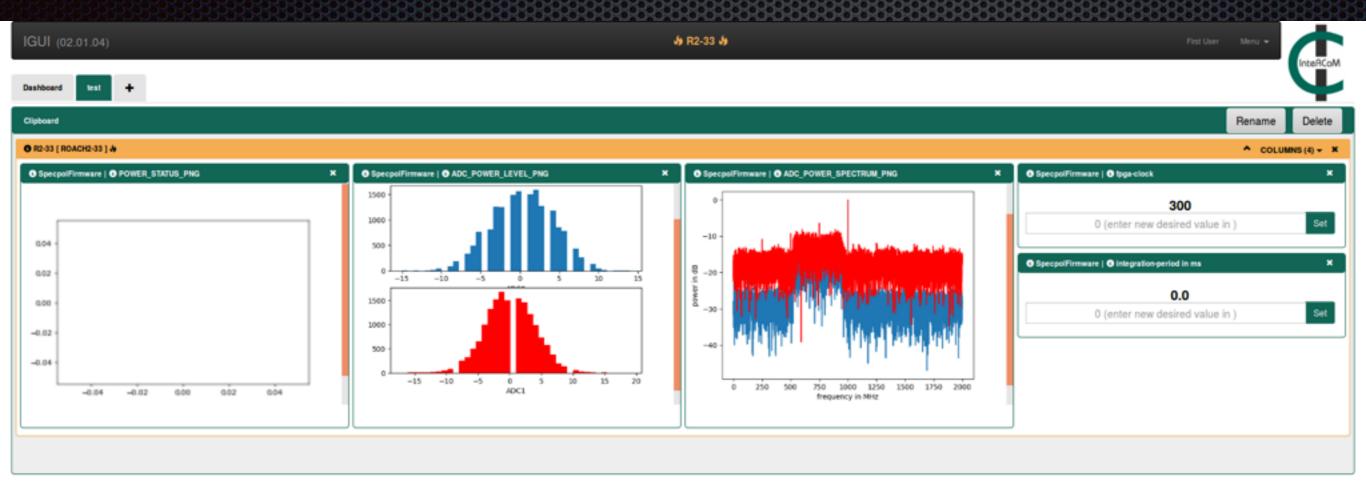
- Must integrate with the telescope control systems and provide real-time feedback for pointing and focus calibration
- Intended as the prototype for TNRT and SKA Prototype dish (slightly different functionality, same framework)

Control Plane

Telescope control system

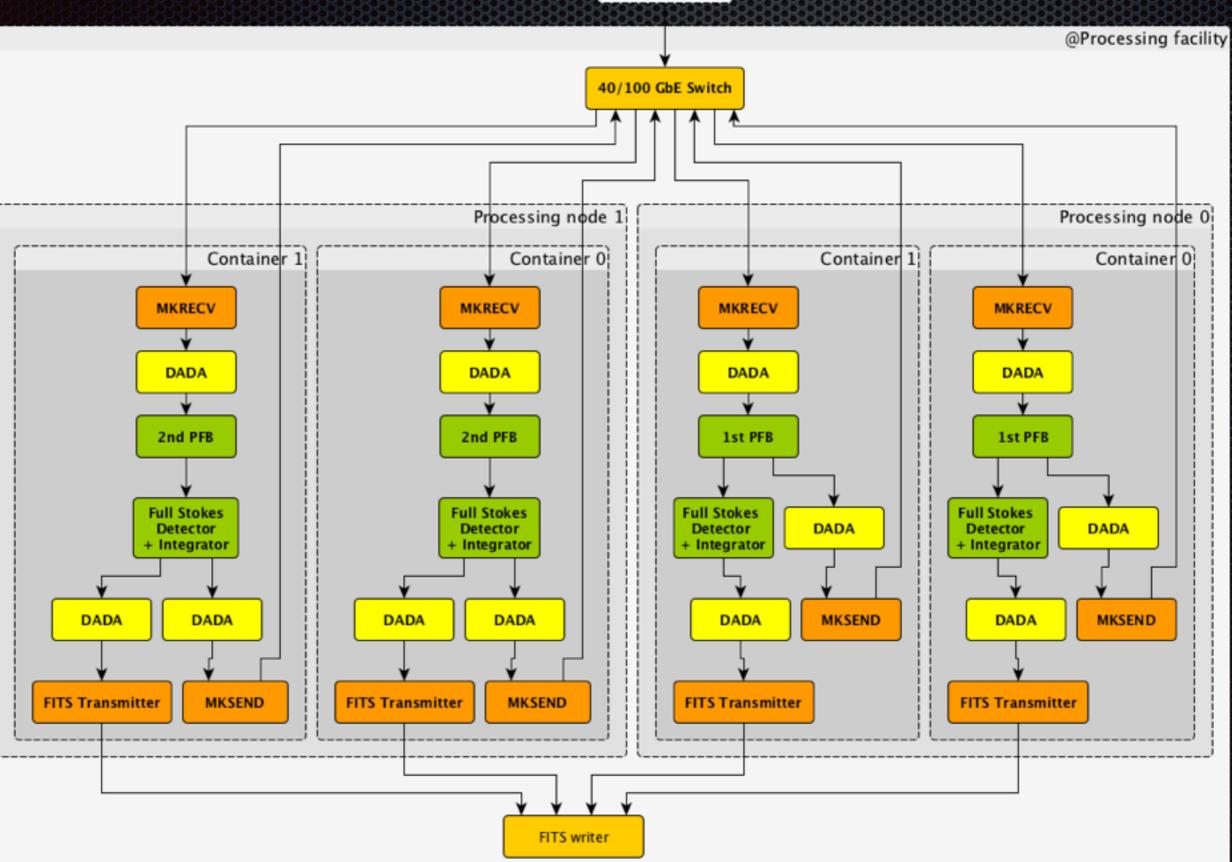


Monitoring interface (iGUI)



ı	ogs											
Frontend-Server-Loga					Receiver-Logs							
	Type ALL	Log Level	Debug Level	# Logs	Apply Filter	Clear Filter	Receiver ALL	ALL	Debug Level	# Logs	- Apply Filter	Clear Filter
	Date	Message			search log(s)		Date	Receiver Mess	191		search log(s)	
l												

Day 1 Functionality



Telescope

Thoughts going forward

- Many risks (noise diodes!), we expect to learn a lot in the coming months
- Future k8s updates will simplify system management across the board
- We are happy to collaborate with any and all. Code is MIT licensed, k8s configs and cluster configs can be made available
- Dependency chains for specific needs should be developed with backpressure deployment
- Need to understand lifetimes of COTS backends better (rolling replacement/upgrade, accelerator changes, etc. etc.)
- Before we needed hardware experts, now we need sys admins!