

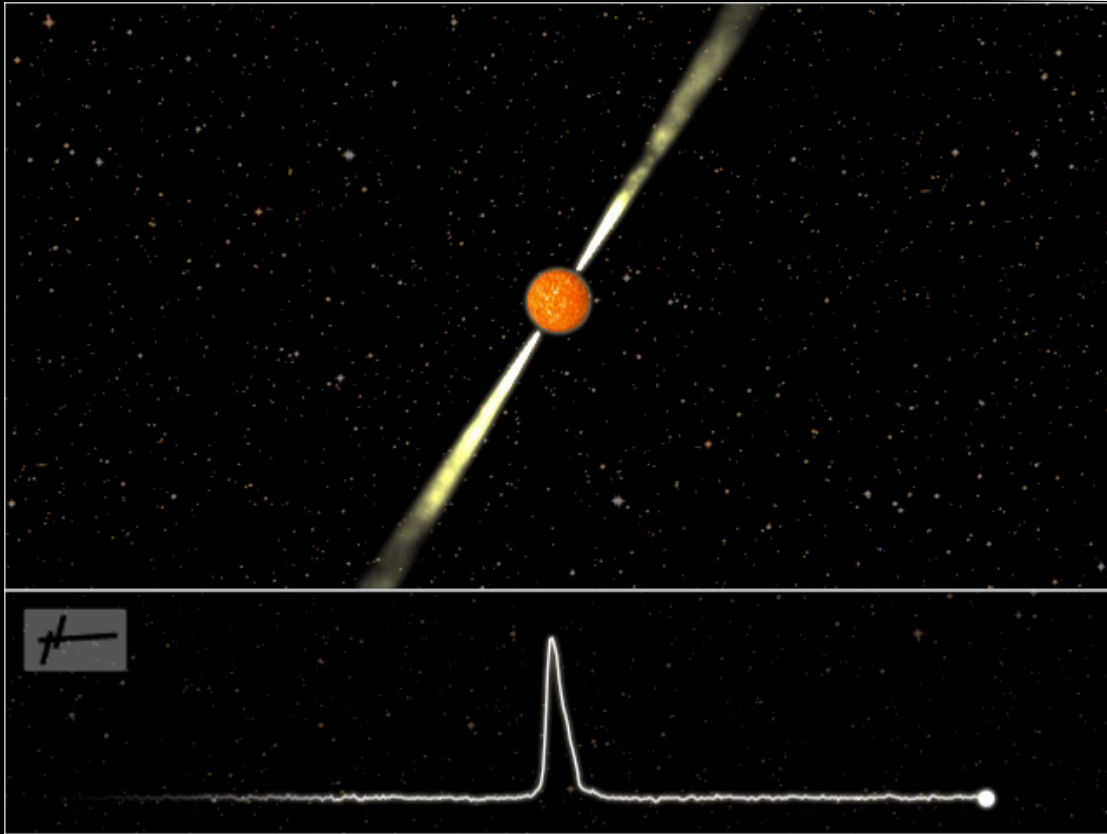
Pulsar Searches and Stuffs

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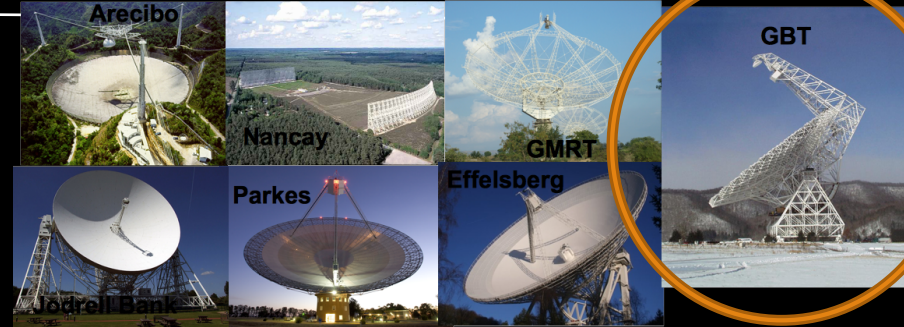
Targeted Pulsar Search (for MSPs)

Credit: <http://www.astron.nl/pulsars/animations>

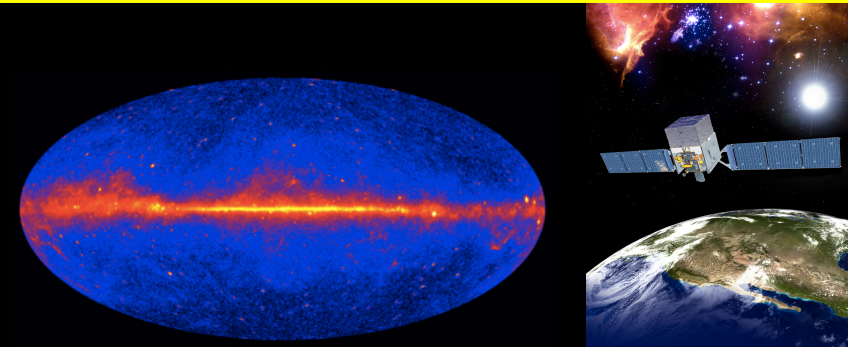


Pulsar is a highly magnetized, rotating neutron star that emits a beam of electromagnetic radiation.

“Physicists’ playground”



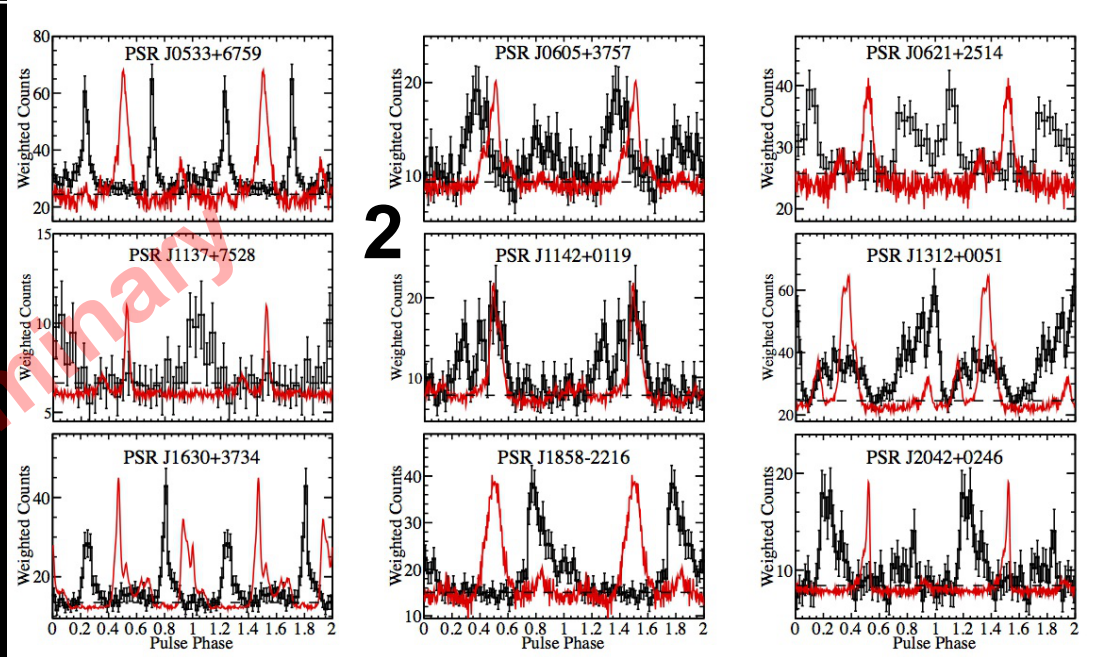
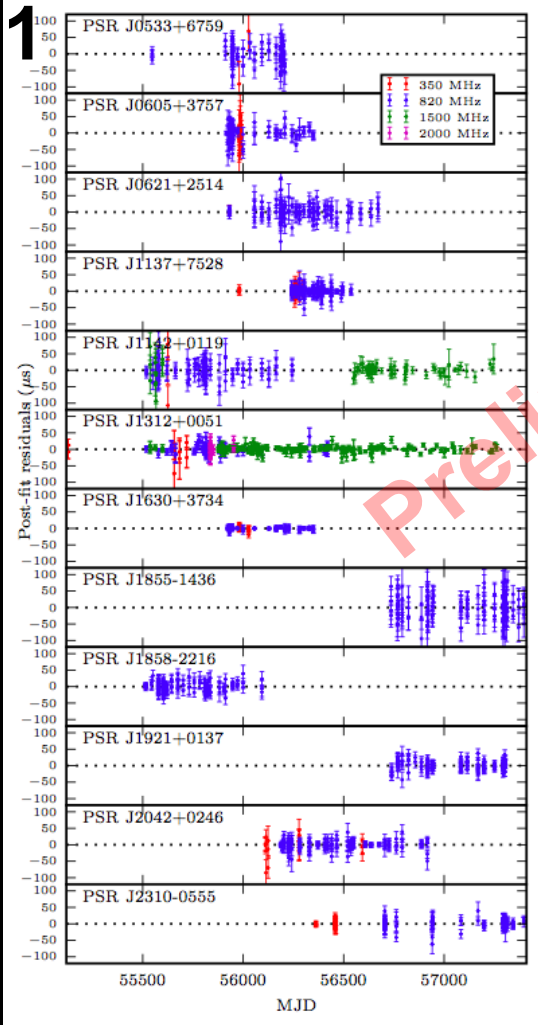
Pulsar is a radio source. The most efficient way to search for new pulsar is to use **single-dish radio telescopes**



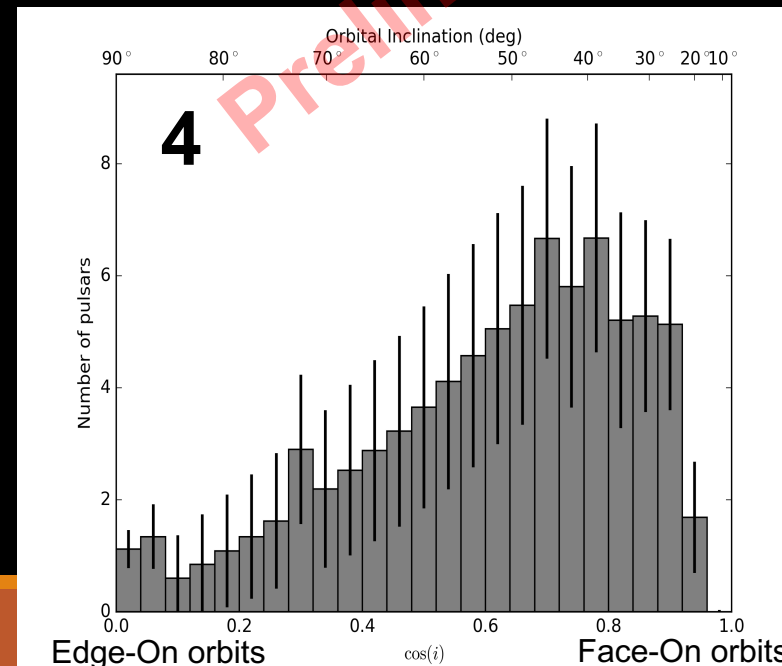
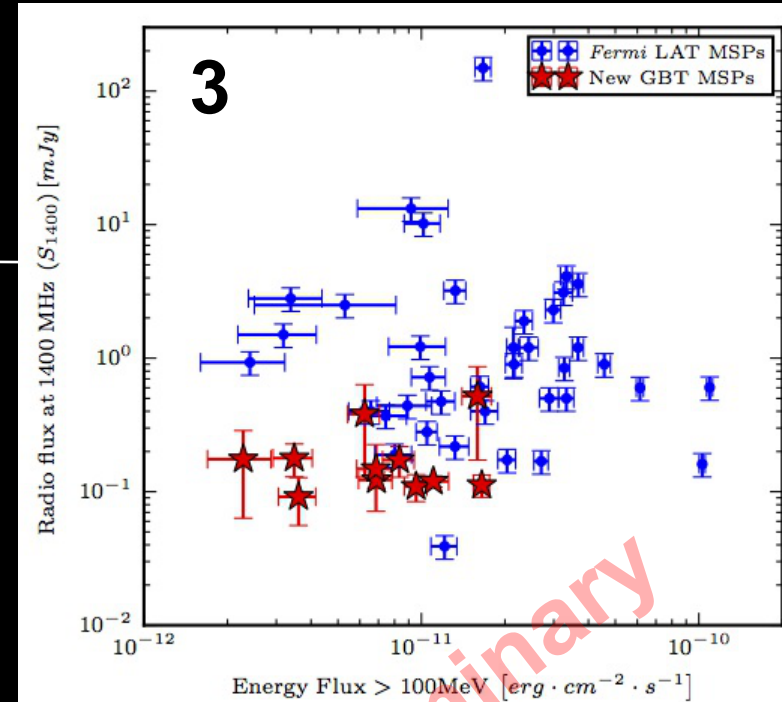
Credit: NASA/DOE/Fermi LAT Collaboration

Millisecond pulsars (MSPs) emit gamma-ray. *Fermi* is the most sensitive gamma-ray telescope; hence, ***Fermi* sources are the best MSPs sources**

Result: "Normal MSPs"



1. Timing solutions for all 12 MSPs (GBT)
2. Every MSPs show gamma-ray pulsations
3. No correlation between radio and gamma-ray flux densities
4. $P_b - M_{\text{com}}$ Analysis: Distribution of $\cos(i)$ is not flat \Rightarrow we are more likely to observe MSPs with face-on orbit



Typeset using L^AT_EX twocolumn style in AASTeX62

Discovery of Twelve Millisecond Pulsars in *Fermi* LAT Unidentified Sources with the GBT telescope

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ABSTRACT

We report the discovery and initial timing solutions of twelve millisecond pulsars (MSPs) from radio searches in the direction of 198 unidentified *Fermi* LAT sources, conducted with the Green Bank Telescope (GBT) at 350 MHz, 820 MHz and 2 GHz. All MSPs (except for the isolated J0533+6759) are in binary systems with likely degenerate He-core white dwarf (He-WD) companions. All twelve of the pulsars are shown to have γ -ray pulsations after folding LAT events using our radio ephemerides. Combining the γ -ray and radio data for all *Fermi* LAT MSPs detected in the radio band, we find that there is no correlation between γ -ray and radio flux densities. We also examine the assumption that the inclination angles (i) of the orbital planes of MSPs with He-WD companions are distributed randomly. Using the orbital period-companion mass relation together with Monte Carlo simulations, we find that the distribution of $\cos i$ of 92 such systems is not flat, but favors higher $\cos i$ (i.e. more face-on orbits).

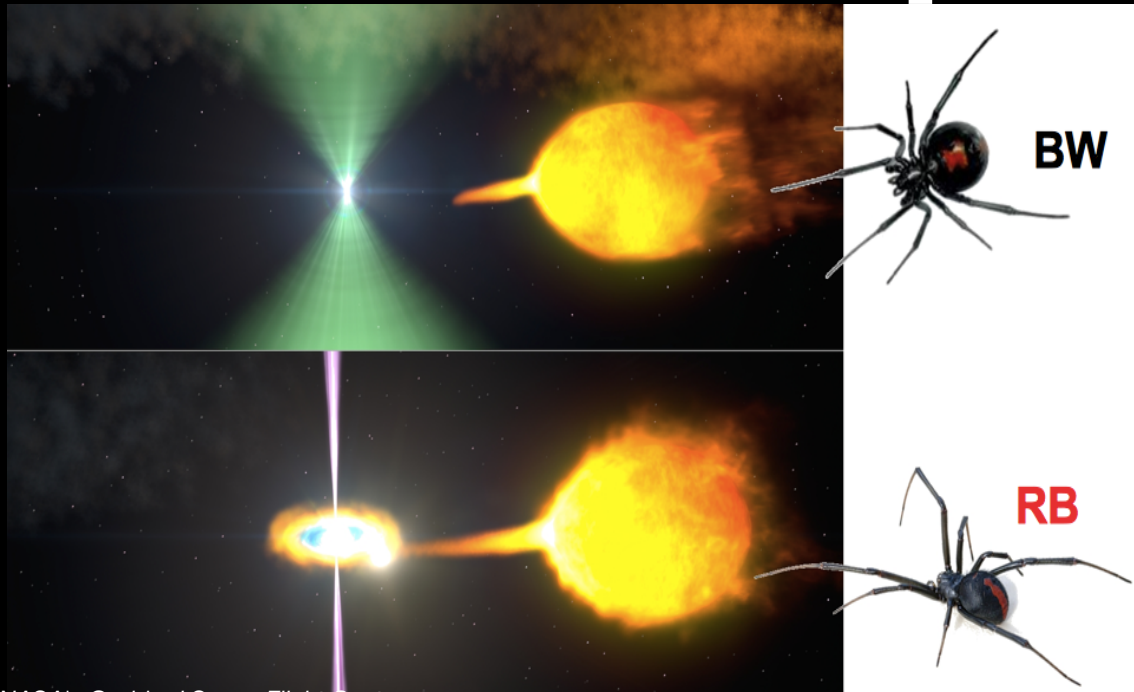
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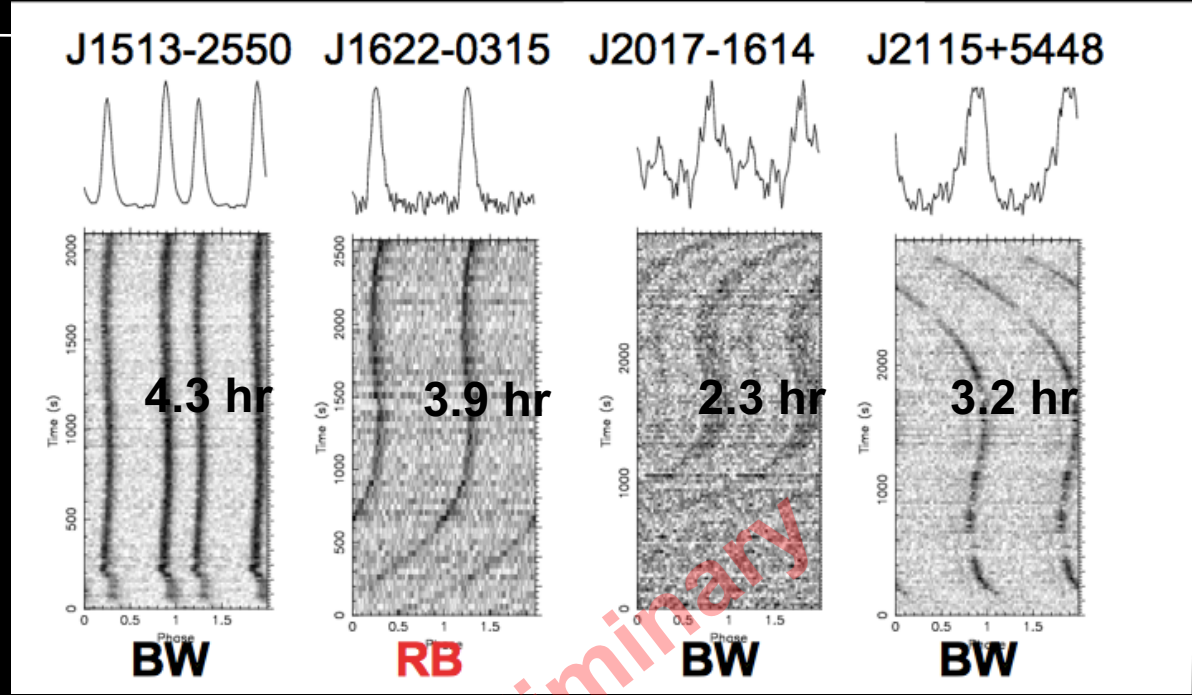
FAST: Pulsar search on *Fermi* sources (Spring 2019)



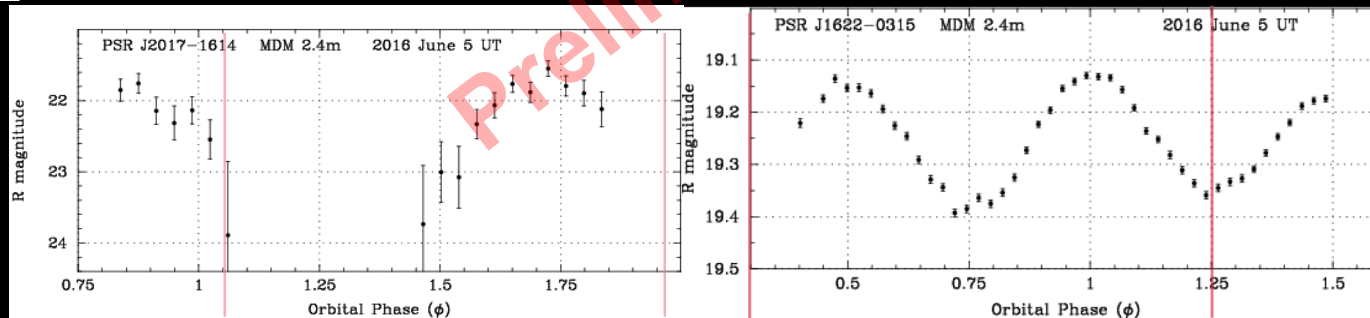
Result: "Spiders"



NASA's Goddard Space Flight Center



- ❑ Spider: A MSP in a compact orbit (orbital period < 1 day).
- ❑ Black Widow (BW): Extremely small companion mass ($\sim 0.01 M_{\odot}$). Ablating away the companion.
- ❑ Redback (RB): Non-degenerated companion. **Accreting mass from the companion.**

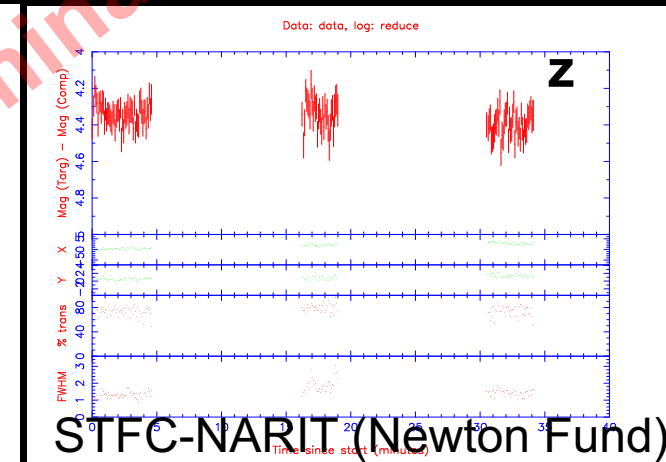
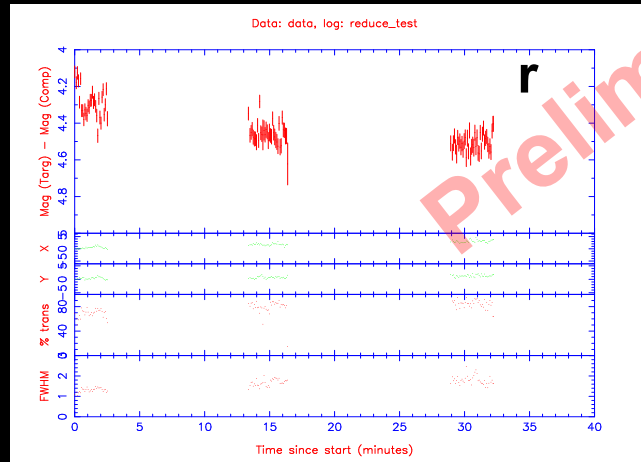
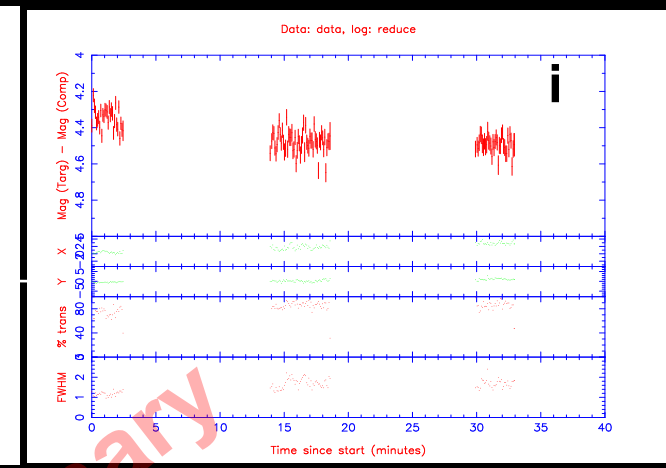
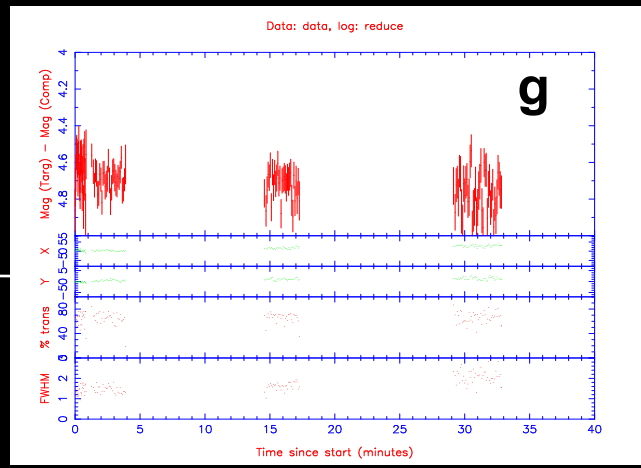
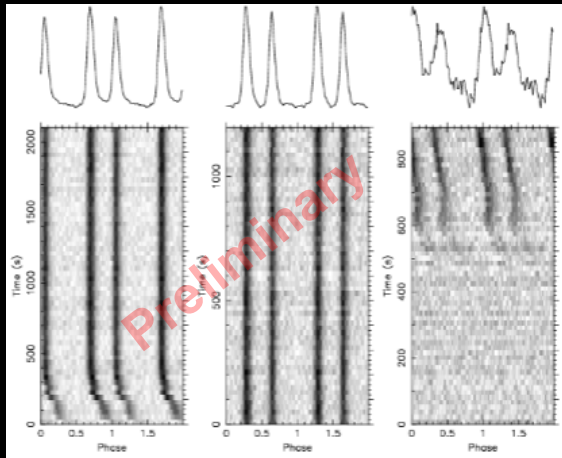
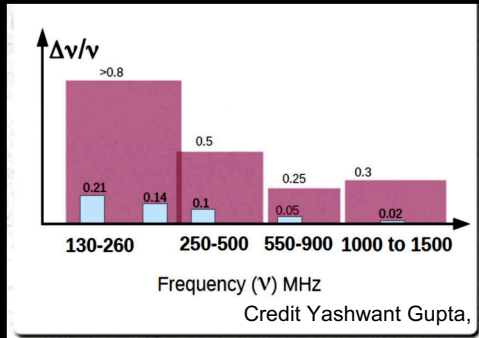


Optical counterparts with the Hiltner 2.4m Telescope

Result: "Spiders"



GMRT vs uGMRT :
Frequency Coverage



STFC-NARIT (Newton Fund)

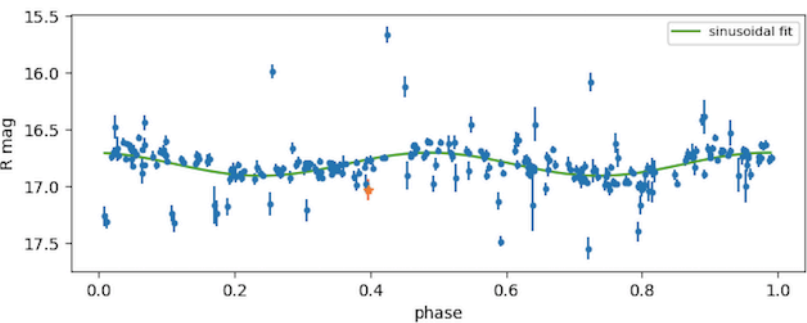
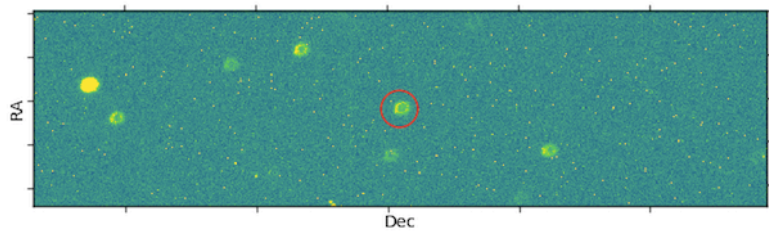
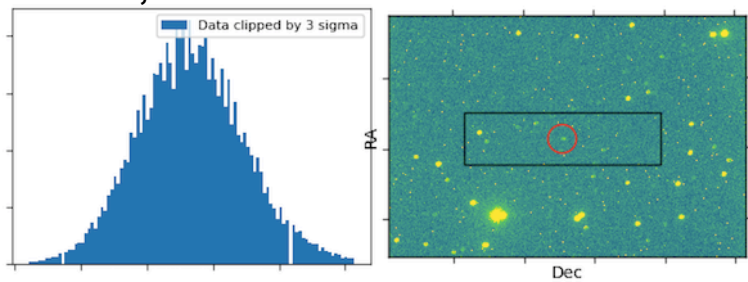
Observing 4 new eclipsing spiders with the uGMRT with multiple Bands

- Long term monitoring PSR J2129-0429 with the 2.4m Thai National Telescope (TNT) using ULTRASPEC
- PSR J2129-0429 is a BW with a flux variability over a long period of time (~10 years) (Breton+2013)
- Data from November 2017 (4 filters) taken by Dr. Puji Irawati

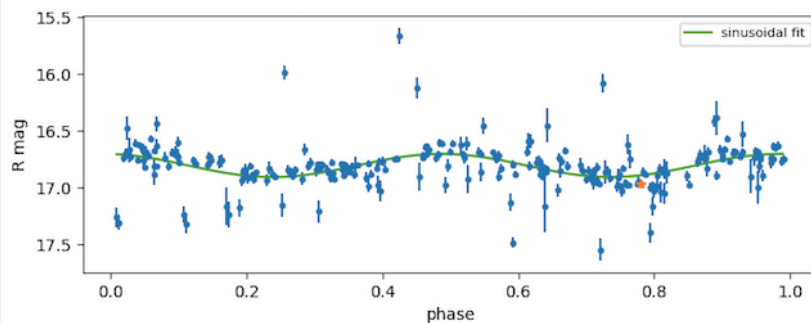
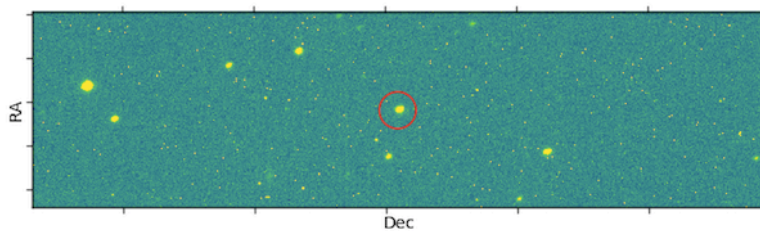
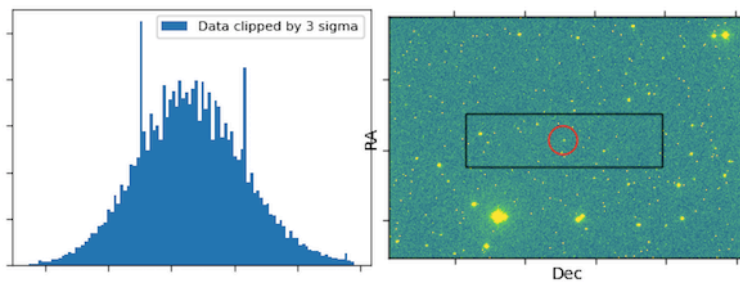


Result: "Spiders": Pt5m data of J2129-0429

PSR J2129-0429 pt5m data
Rene Breton, U. of Manchester
Vik Dhillon, U. of Sheffield



PSR J2129-0429 pt5m data



PSR J2129-0429 pt5m data

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